A B-STAR HIGH LUMINOSITY INDICATOR

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In an effort to sort out BIa supergiants among B-stars for use in delineating spiral structure, a photometric discriminate has been sought in the following way. Wide slit $16\mbox{\normalfont\AA/mm}$ (flux) spectra were obtained of a sample of B stars over a suitable spectral and luminosity range. These spectra were digitized on the PDS Microdensitometer. Numerical filters of various widths were then slid along the spectra and numerous color-indices were formed and tested. An index $\ell_0 = u + v - 1.9$ (1) plotted against $c_{\text{O}}\text{, where }\text{(}$ is a 200Å halfwidth passband centered near 3840 Å and the other magnitudes are on the Strömgren 4-color system, succeeds in separating the Ia supergiants up to Bl. For stars of spectral type Bl and earlier, confusion sets in and the separation is uncertain (although a fair percentage of success is also had for types Bl and BO). This luminosity discriminate is now in use for a large spiral structure program in which all known Be stars (and perhaps eventually all B stars) are being surveyed for Ia supergiants. Spectra for 2-dimensional classification are obtained for the Ia candidates.