78 CONTRIBUTED PAPERS

## REFERENCES

Arquilla, R., and Goldsmith, P.F.: 1985, Astrophys. J. 294, 436. Clark. F.O., and Johnson, D.R.: 1982, Astrophys. J. 263, 160. Milman, A.S.: 1977, Astrophys. J. 211, 128.

## CYANOACETYLENE OBSERVATIONS OF B335

Tatsuhiko Hasegawa, Osamu Kameya, Naomi Hirano, Munezo Seki, and Keiya Takakubo Astronomical Institute, Tohoku University Sendai, Japan

B335 is now recognized as the smallest isolated star forming region. The detection of a Far-IR source and a bipolar flow were succesful, on the other hand, the distribution of the quiet gas is poorly understood. We are trying to determine the density distribution in B335. As the first step, we have carried out  $HC_3N$  (J = 5-4 and 4-3) observations of B335. The observations of the J = 5-4 line have revealed a high density core with a 30"-60" size. The Far-IR source is located just at the center of the core, and the core lies at the center of the bipolar flow. A mean hydrogen molecular density in the core of about  $5\times10^4$  cm<sup>-3</sup> is derived from the line ratio J - 5-4/4-3.

## CO OBSERVATIONS OF A COMETARY GLOBULE IN IC1396

Makoto Nakano, Yoshio Tomita and Hiroshi Ohtani Department of Astronomy, University of Kyoto, Japan Katsuo Ogura Kokugakuin University, Japan Yoshiaki Sofue Nobeyama Radio Observatory, Japan

A cometary globule in IC1396 named "comet tail 6" by Osterbrock (1957), has been observed at CO and  $^{13}$ CO (J = 1-0) lines with a high spatial resolution, 14", with the 45-m radio telescope at the Nobeyama Radio Observatory. The resolution corresponds to a linear size of 0.05 pc at the distance of 750 pc (Matthews 1979). Two possible pre-main sequence stars, LkH $\alpha$  349 and LkH $\alpha$  349/c (Cohen and Kuhi 1979), associated