

The current orbit of Atlas (SXV)

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Abstract. With the success of the Cassini-Huygens mission, the dynamic complexity surrounding natural satellites of Saturn began to be elucidated. New ephemeris could be calculated with a higher level of precision, which made it possible to study in detail the resonant phenomena and, in particular, the 54:53 near mean-motion resonance between Prometheus and Atlas. For this task, we have mapped in details the domains of the resonance with dense sets of initial conditions and distinct ranges of parameters. Our initial goal was to identify possible regions in the phase space of Atlas for which some critical angles, associated with the 54:53 mean motion have a stable libration. Our investigations revealed that there is no possibility for the current Atlas orbital configuration to have any regular behavior since it is in a chaotic region located at the boundary of the 54:53 mean-motion resonance phase space. This result is in accordance with previous works (Cooper et al. 2015; Renner et al. 2016). In this work, we generalize such investigations by showing detailed aspects of the Atlas-Prometheus 54:53 mean-motion resonance, like the extension of the chaotic layers, the thin domain of the center of the 54:53 resonance, the proximity of other neighborhood resonances, among other secondary conclusions. In particular, we have also shown that even in the deep interior of the resonance, it is difficult to map periodic motion of the resonant pair for very long time spans.

Keywords. Resonant dynamics, mean-motion, dynamical maps, Atlas, Prometheus, Saturn.

1. Introduction

The natural satellites of Saturn display a variety of orbital configurations and unique topological features that have intrigued astronomers, physicists and mathematicians for several years (e.g. Peale 1999). Such dynamic environment is responsible for the most diverse gravitational disturbances such that tides, short and long periodic perturbations. In particular, small satellites with their mean radius of the order of ten kilometers or smaller suffer non-negligible quasi-periodic variations in their orbits due to mutual gravitational interactions and, this phenomena imply that many pairs of satellites have their orbits close to commensurability of the mean-motions. In this work, we study the dynamics of Atlas. According to Thomas and Helfenstein (2020), Atlas is a small satellite with radius of 14.9 km orbiting near the outer edge of Ring A, see Table 2.

Spitale et al. (2006) combining data previously obtained by the Voyager spacecraft, the Hubble Space Telescope and ground-based telescopes determined precise ephemeris for these small worlds, in particular for Atlas. In their studies, Spitale et al. (2006) suggested that Atlas orbit is perturbed due to mutual gravitational interaction with Prometheus generating a 54:53 resonant mean-motion and this can be associated with the presence of the chaotic motion.

Table 1. Physical constants for Saturn obtained from Cooper et al. (2015)^a, Horizons^b, J_2 , J_4 and J_6 a dimensionless coefficients of Saturn potential expansion.

Constant	Numerical value
GM^a (Km ³ s ⁻²)	3.7931208×10^7
Equatorial radius ^b (Km)	60268 ± 4
J_2^a	16290.71×10^{-6}
J_4^a	-935.83×10^{-6}
J_6^a	86.14×10^{-6}

Cooper et al. (2015) investigating the existence of the chaotic motion of Atlas suggested by Spitale et al. (2006) by calculating the Fast Lyapunov Indicator (FLI) method to verify the presence of chaos in the Atlas orbit. They analyzed the critical angles associated with the 54:53 resonant mean-motion for a timespan of 20 years. Their integrations showed that the angle associated with the Corotation Eccentric Resonance (ϕ_{CER}) oscillates with temporary oscillation with period about 4.92 years and amplitude around 180° followed by circulation intervals. In contrast, the angle associated with Lindblad Eccentric Resonance (ϕ_{LER}) has temporary oscillation of about 3 years followed by a long period of circulation (see Figure 2 of Cooper et al. 2015).

In an attempt to determine the origin of this alternation between libration and circulation, Renner et al. (2016) studied the Atlas orbit using the CoraLin model proposed by (El Moutamid et al. (2014)). Your results show that the region of the space occupied by the CER resonance is superimposed by the chaotic region.

Following the contributions of Cooper et al. (2015) and Renner et al. (2016), we will investigate the Atlas orbit looking for initial conditions that enable stability for the libration of the ϕ_{CER} and ϕ_{LER} angles. In additions, we identify possible causes for instability described in Cooper et al. (2015).

2. Methodology

In this work, we follow the methodology given in Callegari and Yokoyama (2010, 2020) and Callegari et al. (2021) and numerically integrate the exact equations of the motion for a system of N satellites mutually disturbed orbiting under the action of the main terms of the Saturn's potential expanded up to second order. In our simulations, for brevity we show the main results considering only $N = 2$, that is, a system formed by Atlas and Prometheus. The justification for this choice stems from Atlas orbital elements can well be determined over the influence of Prometheus, without considering the other effects arising from the perturbation caused by Pandora or other companion satellite (Spitale et al. 2006; Cooper et al. 2015; Renner et al. 2016). It is worth to note that we have considering more general models in this work, but for brevity only the result within the domains of the three-body problem are shown here.

Two different models were used: i) the system of exact differential equation (Eq. 1–5) present in Callegari and Yokoyama (2010), where the equations of motion are integrated under the influence of the terms J_2 and J_4 and; ii) direct application of the Mercury package (Chambers 1999) with the addition of the term J_6 . In both cases i) and ii) we apply the Everhart code “RA15” to solve systems of ordinary differential equations (Everhart 1985).

Physical parameters for Saturn and initial conditions for Atlas and Prometheus provided by the Horizons system of ephemerides (<http://sdd.jpl.nasa.gov/horizons.cgi>) are listed in Tables 1 and 2, respectively. The masses were obtained from Thomas and Helfenstein (2020) and the details for the numerical simulations are described in the caption of the respective figure.

Table 2. Mass and osculating orbital elements at Epoch 2000 January 1 00: 00: 00.0000 UTC computed from Ephemeris system - *Horizons*, in December 13, 2019.

	<i>Mass (Kg)</i>	<i>a (Km)</i>	<i>e</i>	<i>I (°)</i>	<i>ω (°)</i>	<i>Ω (°)</i>	<i>n (°/day)</i>
Atlas	5.75×10^{15}	138325.32	0.00591	0.00419	200.7	235.45	592.63
Prometheus	1.5×10^{16}	140246.44	0.00252	0.00801	201.36	309.14	581.87

Often, some oscillating elements obtained through numerical integration can be highly influenced by the term J_2 , causing a fast frequency component in the pericenter (ω), a fact that makes difficult its interpretation (Greenberg 1981). Thus, as an additional effort, we calculate the geometric elements through the direct application of the algorithm described by Renner and Sicardy (2006).

3. The current orbit of Atlas and the 54:53 Prometheus-Atlas mean-motion

Atlas finds itself orbiting a region of dynamic complexity. The understanding of the 54:53 resonant mean-motion make it necessary to determine the stability of its orbit.

3.1. The critical angles ϕ_{CER} and ϕ_{LER}

Resonant phenomena are common among satellites of Saturn, in particular the Corotation Eccentricity Resonance (CER) and Lindblad Eccentricity Resonance (LER). According to Callegari et al. (2021) the CER resonance occurs when the conjunction between disturber (larger body) and particle (smaller body) happens near the pericenter of the disturbed one, on the other hand, the LER resonance occurs when the conjunction occurs near the pericenter of the particle. The knowledge of the physical behavior of these two angles makes it possible to predict where the conjunctions between the two bodies occur, in addition to verifying the orbital stability of the system. For the 54:53 resonant mean-motion between Prometheus and Atlas, the angles ϕ_{CER} and ϕ_{LER} can be written as an angular combination of the form $\phi_{CER} = 54\lambda_P - 53\lambda_A - \varpi_P$ and $\phi_{LER} = 54\lambda_P - 53\lambda_A - \varpi_A$, where $\lambda = \varpi + l$ represents the mean longitude, ϖ the pericenter longitude, l the mean anomaly, whereas A and P represent Atlas and Prometheus, respectively. Physically, the libration of the ϕ_{CER} angle around 180° means that the conjunctions occur close to the apocenter of Prometheus. In the case of the current orbit of Atlas, Fig. 1 shows that there are periods of alternation between oscillation and circulation of ϕ_{CER} .

On the other hand, the ϕ_{LER} angle has a brief period of oscillation, about 5,000 days, followed by a long period of circulation, as shown in Fig. 1.

Cooper et al. (2015) state that Atlas could be added to the list of satellites that have both critical angles librating, however, for a satellite similar to Atlas, we cannot make such an affirmation, as we do not obtain stability for the dynamics of these angles as it occurs for Anthe (Callegari and Yokoyama 2020) and Methone (Callegari et al. 2021). See also El (El Moutamid et al. (2014).

Expanding the investigation done by Cooper et al. (2015) we will look for a possible orbital configuration that can guarantee the stability of the libration of ϕ_{CER} and ϕ_{LER} . For this task, we will use the phase space mapping through Fourier spectra, a methodology described in Callegari and Yokoyama (2010, 2020) and Callegari et al. (2021).

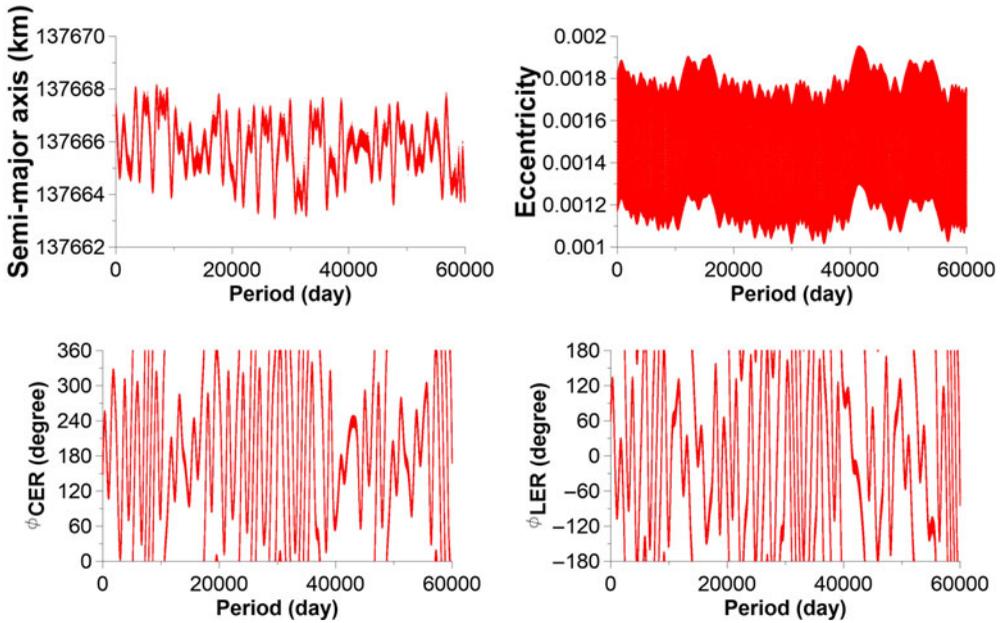


Figure 1. Simulation performed with the initial conditions given in Table 2 for an Atlas-like satellite. Semi-major axis and eccentricity geometric elements and critical angles in geometric elements associated with CER and LER type resonances. Note the lack of stability for the oscillation of ϕ_{CER} and ϕ_{LER} . The equations were integrated for 60,000 days with a 0.06 day step using the Mercury package.

3.2. The mapping of the 54:53 Prometheus-Atlas resonant mean-motion

We will now explore the resonant phase space domain. The mapping is carried out considering the domain of frequencies obtained with the Fourier spectrum for a set of numerically integrated orbits, whose initial conditions are close to the real orbit of Atlas.

Callegari and Yokoyama (2010, 2020) and Callegari et al. (2021) analyze the phase space around the orbital neighborhood of a satellite according to the value of the spectral number N . For the construction of this N it is considered an established value, usually 5% of the reference amplitude for the Fourier spectrum for each of the individual orbits of these test satellites. Then, N provides the number of peaks in the spectrum that are greater than or equal to the pre-set reference value.

We will now investigate the dependence of the resonant motion with the semi-major axis and the eccentricity for an Atlas-like satellite. Fig. 2 shows two dynamic maps built on a dense grid of initial conditions for (a_0, e_0) . The red star represents the initial condition $(a_0, e_0) = (138325.32 \text{ km}, 0.0059)$ for Atlas on the initial date (see Table 2).

The map at the left in Fig. 2 represents the phase space of the spectral domain for the semi-major axis and the map on the right represents the spectral domain for the orbital eccentricity. We can identify four distinct regions:

- a) A light “blue eye” region, with N less than 30, in the range $[138321 \text{ km}, 138326 \text{ km}] \times [0.004, 0.006]$ indicated by **C**. In this region all angles ϕ_{CER} are oscillating around 180° but the alternation occurs between the oscillation and circulation. However, for the initial conditions $(138322 \text{ km}, 0.004)$ blue disc and $(138323 \text{ km}, 0.005)$ magenta disc, in Fig. 2, the angle ϕ_{CER} is librating, see Fig. 3 items (b), (e) and (f). This would probably be the region of the corotation zone associated to the 54:53 Prometheus-Atlas mean-motion resonance, but, the absence of stability

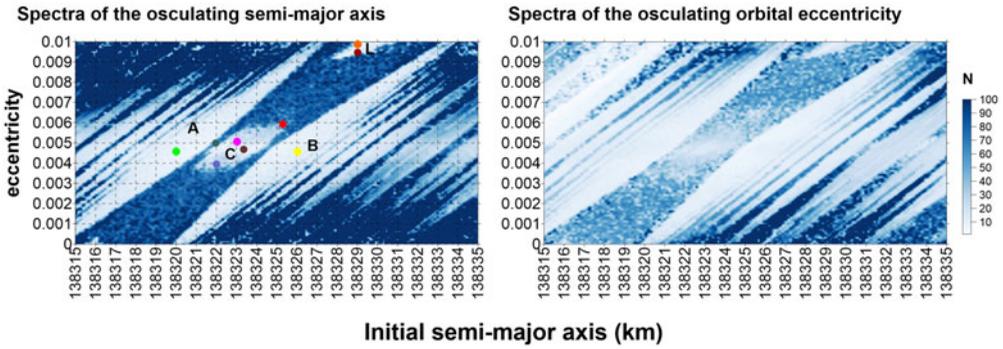


Figure 2. Dynamic map constructed from the spectrum of the osculating semi-major axis (left) and orbital eccentricity (right) of 15,000 Atlas-like satellites (numerical scheme i). N is the spectral number with 5% reference amplitude. C and L are the regions in which ϕ_{CER} and ϕ_{LER} angles have periods of oscillation of approximately 30,000 days and 15,000, respectively, but alternating their behavior. For this simulation, Prometheus and Atlas were used, the set of differential equations had been integrated for 258 years with an interval of 0.18 days. Initial conditions are given in Table 2. The initial element of the real Atlas are indicated by a red star. Different full discs correspond to initial conditions (a_0, e_0) of the orbits shown in Figure 4.

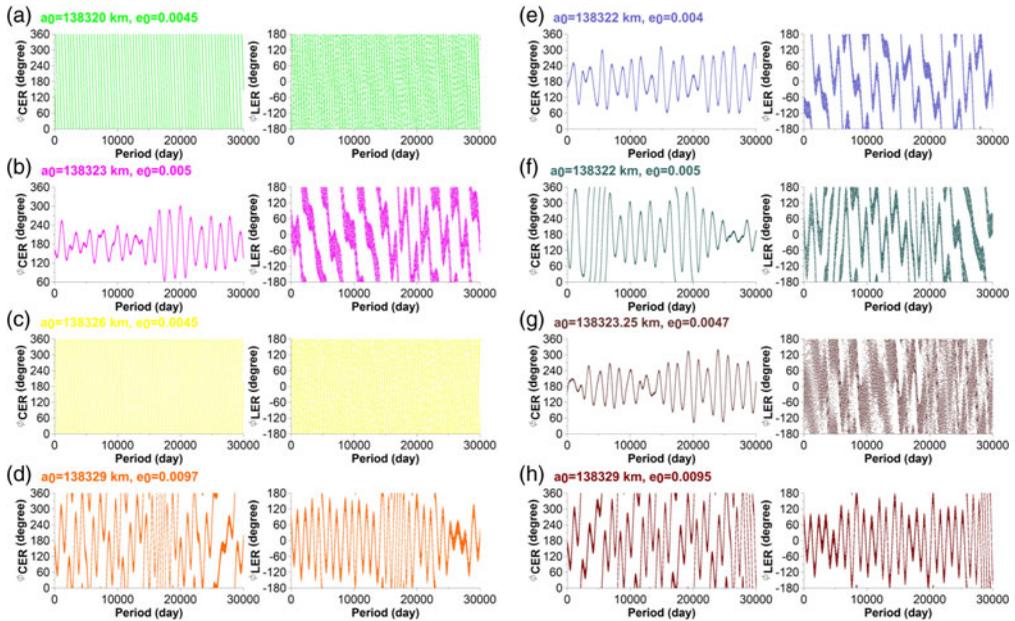


Figure 3. Geometries angle ϕ_{CER} and ϕ_{LER} obtained with the numerical simulations (numerical scheme ii) corresponding to the colored discs in Fig. 2. In a) and c) both are circulating; b) e) and g) the angle ϕ_{CER} are librating and ϕ_{LER} circulating; and g) alternation between libration and circulation of angle ϕ_{CER} and circulation of angle ϕ_{LER} ; d) and h) libration of angle ϕ_{LER} and circulation of angle ϕ_{CER} .

in the libration for the other initial conditions, Fig. 3 (f), does not allow this region to be considered as such.

- b) The dark blue region, which fills practically the entire map, where the spectrum of the test satellites has a large value for N . In general, such behavior can be associated with chaotic orbits and strong or irregular disturbances in their motions (Callegari and Yokoyama 2020).
- c) Region between $[138327 \text{ km}, 138330 \text{ km}] \times [0.009, 0.01]$, indicated by **L**. Location where the angle ϕ_{LER} shows oscillation around 0° . The absence of stability in the libration does not allow this region to be considered as such, see Fig. 3 items (d) and (g).
- d) Bands A and B: in this region angles ϕ_{CER} and ϕ_{LER} show themselves circulating in a retrograde direction for a brief period of time, about 500 days (band A) and for region B, the angles are circulating in a prograde direction with a period of approximately 400 days. See Fig. 3 items (a) and (c).

We can verify, with the help of the dynamic map given in Figure 2, that there is no region with (a_0, e_0) in which there is stability for the critical angles ϕ_{CER} and ϕ_{LER} libration, neither for both critical angles.

We can observe that for the initial conditions given in Table 2, Atlas is found close to the edge of a possible chaotic region and this is probably responsible for the alternations observed in the angles ϕ_{CER} and ϕ_{LER} observed in Cooper et al. (2015) (See Figure 1).

(El Moutamid et al. (2014) have applied a general development of corotation and Lindblad resonances (CoraLin model), ideal to be analyzed with the surface of section technique. The y-axes of the section are given by a variable χ , representing the resonance width (see Table 1 in (El Moutamid et al. (2014)) which is plotted against the corotation angle ϕ_{CER} .

In Fig. 4, we show several plots of $\chi \times \phi_{CER}$ for some orbits obtained numerically through the Mercury package taken in the regions (a), (c) and (d), where we consider the full values of χ obtained from numerical orbits, not the sections. Also we utilize $a' = 137665.519 \text{ km}$ as the reference geometric center of the corotation zone (Renner et al. 2016).

We have the following conclusions:

- A) When values for the semi-major axis are between 138322 and 138323 km, χ assumes small values while ϕ_{CER} alternates between libration and circulation. See Figure 5 in Renner et al. (2016) and Figures 3 and 4 of this work.
- B) For large values of eccentricity, χ becomes diffuse in the $\chi \times \phi_{CER}$ plane, while ϕ_{CER} is circulating.

4. Final considerations

The use of the dynamic map obtained through the spectrum of individual orbits revealed that:

- i) Atlas is located on the edge of a possible chaotic region. Physically, this location implies the absence of stability for critical angles ϕ_{CER} and ϕ_{LER} found by Cooper et al. (2015);
- ii) for the interval of semi-major axis and eccentricity values given by $[138322 \text{ km}, 138323 \text{ km}] \times [0.004, 0.005]$ the map reveal an initial conditions (a_0, e_0) in which we can find a stable libration for ϕ_{CER} and also, initial conditions in which the angle ϕ_{CER} presents episodes of libration oscillation followed by circulation;

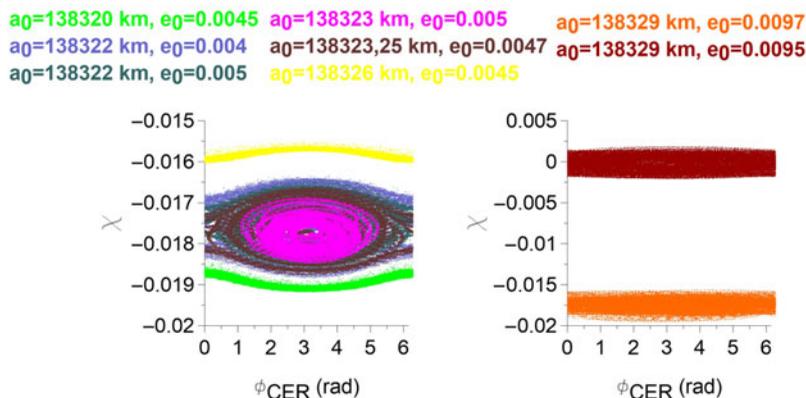


Figure 4. Plot for $\chi \times \phi_{CER}$ of several orbits with initial conditions gives in coloured full discs in Fig. 2.

- iii) for eccentricity values greater than 0.009, and semi-major axis between 138328 and 138330 km we identified the region in which periods of libration occur for the angle ϕ_{LER} , as well as in ii) we did not find a region our intial condition in which there is stable libration for ϕ_{LER} .
- iv) for regions A and B in Fig. 2, the orbits were presented as circulating according to (El Moutamid et al. (2014)). Such feature indicates that the Atlas-like particle is far from a' .

Therefore, we conclude that Atlas is not deeply inserted in to the 54:53 resonance, due to the lack of stability in the ϕ_{CER} and ϕ_{LER} angle libration. This fact implies that Atlas cannot be added to Saturn's list of natural satellites that have both critical angles librating.

References

- Callegari Jr., N., Yokoyama, T. Numerical exploration of resonant dynamics in the system of Saturnian inner Satellites. *Planetary and Space Science* 58, 1906–1921 (2010).
- Callegari Jr., N., Yokoyama, T. Dynamics of the 11:10 Corotation and Lindblad Resonances with Mimas, and Application to Anthe, Icarus, 348, 113820 (2020).
- Callegari Jr., N., Rodrigues, A., Ceccatto, D. T. The current orbit of Methone (S/2004). *Celest Mech Dyn Astr*, accepted for publication (2021).
- Chambers, J. E. A hybrid symplectic integrator that permits close encounters between massive bodies. *Montly Notices of the Royal Astron. Society* 304, 793–799 (1999).
- Cooper, N. J. Renner, S. Murray C. D., Evans M. W. Saturn's inner satellites orbits, and the chaotic motion of Atlas from new Cassini imaging observations. *The Astronomical Journal* 149, 27–45, (2015).
- El Moutamid, M. Renner, S. Sicardy, B. Coupling between corotation and Lindblad resonances in the elliptic planar three-body problem. *Celest Mech Dyn Astr*, 118, (2014).
- Everhart, E. An efficient integrator that uses Gauss-Radau spacings. In: *IAU Coloquium* 83, 185–202 (1985).
- Greenberg, R. Apsidal Precession of Orbits about an Oblate Planet. *The Astronomical Journal* 86, 912–914 (1981).
- Murray, C. D, Dermott, S. F. *Solar System Dynamics*, Cambridge University Press (1999).
- Peale, S. J. Origin and Evolution of the Natural Satellites. *Annual Review of Astron. and Astrophys.* 37, 533–602 (1999).
- Renner, S., Sicardy, B. Use of the Geometric Elements in Numerical Simulations. *Celestial Mechanics and Dynamical Astronomy* 94, 237–248 (2006).

- Renner S. Cooper J. N. El Moutmaid M.; Sicardy B. Vienne A.; Morray C. D. Sarlenfest M. Origin of the chaotic motion of the Saturnian satellite Atlas. *The Astronomical Journal*, 151 (2016).
- Spitale, J. N., Jacobson, R. A., Porco, C. C., Owen, Jr, W. M. The Orbits of Saturn's Small Satellites Derived from Combined Historic and Cassini Imaging Observations. *The Astronomical Journal* 132, 792–810 (2006).
- Thomas, P. C., Helfenstein, P. The small inner satellites of Saturn: Shapes, structures and some implications. *Icarus* 344, 113355 (2020).