MID-INFRARED SPECTROSCOPY OF FOUR 21µm EMISSION BAND SOURCES

K. JUSTTANONT and M. J. BARLOW

Dept. of Physics & Astronomy, University College London, Gower Street, London WC1E 6BT, U.K.

and

C. J. SKINNER

Institute of Geophysics & Planetary Physics, L-413, Lawrence Livermore National Laboratory, P.O.Box 808, Livermore, CA94551-9900, U.S.A.

We report 10 and 20 μ m spectroscopic observations of four C-rich post-AGB objects which exhibit the unidentified emission feature at 21 μ m. The observations were carried out in October 1990 and May 1991 using CGS3 on UKIRT. The spectral resolutions were 70 for the wavelength range of 7.4–13.3 μ m and 80 for the region between 15.4–24.1 μ m. Three of the sources reported here are from the list of Kwok, Volk & Hrivnak (1989), i.e., IRAS 04296+3429; IRAS 07134+1005 and IRAS 22272+5435. Figure 1 shows the full spectrum of IRAS 04296+3429 and IRAS 22272+5435. The 10 μ m spectra of these objects exhibit UIR bands whose peaks all fall longwards of the usual peak wavelengths associated with such features. This may be related to the fact that they are the lowest excitation objects so far found to exhibit UIR emission bands. We also found narrow emission features superimposed on the long wavelength wing of the 21 μ m emission bands of IRAS 04296+3429 and IRAS 22272+5435. The fourth object we observed, SAO 163075, was found to also exhibit a (weak) 21 μ m emission feature. However, there is no PAH features in the 10 μ m region, apart from the plateau at 12 μ m.

Reference

Kwok, S., Volk, K.M., Hrivnak, B.J., 1989, ApJ, 345, L51

