Statistical characteristics of horizontal proper motions in the vicinity of pores

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Abstract. Movement and coalescence of magnetic elements could explain the evolution and growth of pores. There have been numerous studies focusing on flow fields in and around individual pores. We have undertaken a systematic study of the statistical properties of such flows. Data of the *Hinode* Solar Optical Telescope offer an opportunity for this type of research, because of the uniform data quality and absence of seeing so that pores can directly be compared in different environments and at various stages of their evolution. We analyzed about 220 timeseries of G-band images using local correlation tracking. The thus computed flow maps make up a database, which covers various scenes on the solar surface. We use an isolated pore to illustrate the statistical parameters collected for further statistical analyses, which include information about morphology, horizontal flows, evolutionary stage (young, mature, or decaying), complexity of the surrounding magnetic field, and proximity to sunspots or cluster of G-band bright points.

Keywords. Sun: magnetic fields, Sun: photosphere, Sun: sunspots, methods: data analysis

1. Introduction

Photospheric flow fields play an important role in the formation, evolution, and decay of solar magnetic features. Our Local Correlation Tracking (LCT) algorithm (Verma & Denker 2011) is based on ideas put forward by November & Simon (1988). The algorithm was adapted to subimages with sizes of 32×32 pixels (2560 km \times 2560 km), so that structures with dimensions smaller than a granule contribute to the correlation signal. Because cross-correlation techniques are sensitive to strong intensity gradients, a high-pass filter was applied to the entire image, suppressing gradients related to structures larger than granules. The high-pass filter is implemented as a Gaussian with a FWHM of 15 pixels (1200 km).

2. Statistical properties of photospheric flow fields around pores

Starting point for identifying pores is a time averaged (one-hour) G-band image, which has been corrected for limb darkening. Some additional smoothing and subsequent thresholding at 90% of the quiet-Sun intensity I_0 yields a binary mask marking all dark features in the image. The intensity threshold roughly corresponds to the transition from penumbra to umbra in a sunspot. Morphological closing smoothes boundaries and eliminates isolated dark pixels. Only features with a size between 1 Mm² and 225 Mm² are kept. These are conservative limits for magnetic knots and sunspots, respectively. The next criterion is that a pore must have a dark core, i.e., pixels with an intensity lower than $0.7I_0$ should cover at least 0.5 Mm². Nevertheless, rudimentary or orphan penumbrae are sometimes erroneously classified as pores. However, in almost all cases, these features are larger than 20 Mm², have an intensity contrast lower than 20%, and possess a high positive kurtosis. Thus, these feature can also be eliminated from the binary mask.



Figure 1. Average (60-minute) G-band image (*left*) of an isolated pore observed on 2007 January 17. The arrows indicate magnitude and direction of the horizontal proper motions. An arrow with the length of the grid spacing indicates a speed of 0.5 km s⁻¹. Azimuthally averaged profiles (*right*) of the G-band intensity (*dotted*), the time-averaged G-band intensity (*solid*), the horizontal flow velocities (*dash-dotted*), and the divergence (*long-dashed*). The *dashed* curve is the G-band intensity profile after convolving with a Gaussian kernel (*gray*, FWHM = 1.2 Mm).

Statistical properties of pores and flows can now be computed for each contiguous region in the binary mask. Deriving average values for the flows is very easy because the time-averaged G-band image exactly matches the LCT flow maps. Standard tools for blob analysis (Fanning 2003) are used to derive parameters describing the morphology of pores and the associated flow field. We measured the area $A = 3.0 \text{ Mm}^2$ covered by the pore, the length L = 7.2 Mm of the circumference, the ratio A/L = 0.41 Mm, the minimum $I_{\rm min} = 0.39$ and the mean $I_{\rm mean} = 0.61$ intensities, the average flow speed $v_{\text{mean}} = 0.16 \text{ km s}^{-1}$, the divergence $\nabla v = -1.8 \times 10^{-4} \text{ s}^{-1}$, and the vorticity $\nabla \times v = -1.1 \times 10^{-6} \text{ s}^{-1}$. Fitting an ellipse to each pore yields among other parameters the eccentricity e = 0.47 and the angle $\alpha = 62.1^{\circ}$ of the major axis with the horizontal (all values refer to the pore shown in Fig. 1). Radial averages beyond the pore's boundary are only computed, if there are segments covering at least 180° in azimuth which are free of any other dark features. An example of blob analysis and azimuthally averaged profiles is presented in Fig. 1. A striking feature in the intensity map is a bright ring around the pore and an annular structure of positive divergence, which envelops the bright intensity ring like an onion peel. The respective divergence centers in the annular structure are related to exploding granules. Inflows inside the pore and outflows in the exterior are commonly found in pores, and the radial velocity profile has a slow gradient inside pore. The detailed results of the statistical study are deferred to an upcoming publication.

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