INTERACTION BETWEEN MAGNETIC FIELDS AND MOLECULAR CLOUDS ON THE RADIO ARC

M. TSUBOI, T. HANDA[†], M. INOUE, J. INATANI, and N. UKITA Nobeyama Radio Observatory, Minamisaku, Nagano 384-13, Japan [†]Institute of Astronomy, University of Tokyo, Tokyo 181, Japan

We have made observations of the Galactic center region in the CS J=1-0 and J=2-1 lines with the 45-m telescope at Nobeyama Radio Observatory in order to depict the fine structure and the activity of this region. We propose that several expanding molecular structures in this region are originated by successive supernova explosions at least of 10^{2-3} times within 10^6 years. These may be responsible for strong magnetic fields in the Galactic center region. The interaction between one of the expanding molecular structures and the "vertical" filaments is found. From the

200 -45 V -15 m/s

-45 V -15 m/s

150

VERTICAL FILAMENT

100 80 60 40 20

appearance of the interaction, the magnetic field in the "vertical" filaments is estimated to be about 3-5 mGauss.

A molecular filament crosses the "vertical" filaments at the compact polarized source at 10 GHz. Molecular components, flowed from the crossing points, may sweep ionized matter off the gap between the vertical filaments and make the compact polarized source (see Fig. 1).

Figure 1. Integrated molecular emission of CS J=2-1 from -45 to -15 km/s at the compact polarized source at 10 GHz (Tsuboi et al., 1985, 1986; Seiradakis et al., 1985). The origin of coordinates is $\alpha = 17^{\rm h}43^{\rm m}16^{\rm s}$, $\delta = -28^{\circ}51'15''$. The contour interval is $\int Ta^{\rm t}dv = 0.3$ K·km/s. Two dotted lines show the positions of the "vertical" filaments.

References

Seiradakis, J.H. et al. (1985) Nature 317, 697.

Tsuboi, M. et al. (1985) Publ. Astron. Soc. Japan 37, 359.

Tsuboi, M. et al. (1986) Astron. J. 92, 818.

RIGHT ASCENSION OFFSET(")

38

R. Beck et al. (eds.), Galactic and Intergalactic Magnetic Fields, 381. © 1990 IAU. Printed in the Netherlands.