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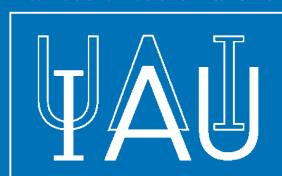
# Neutron Stars and Pulsars: Challenges and Opportunities after 80 years

*Edited by*

Joeri van Leeuwen

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NEUTRON STARS AND PULSARS:  
CHALLENGES AND OPPORTUNITIES AFTER 80 YEARS

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*COVER ILLUSTRATION:*

1932 – 2012:

DISCOVERY OF THE NEUTRON – A NEUTRON STAR IN A SUPERNOVA

This diptych combines the 1932 detection of the neutron with the state of modern neutron-star and pulsar research, in 2012.

In the left-hand side photograph, neutrons have collided with the atoms in a layer of paraffin wax, ejecting a proton. The proton path is visible in the ionization chamber.

The right-hand panel shows an optical (HST) and X-ray (Chandra) false-color image of supernova remnant 1E 0102-7219. Overlaid for illustration is radio (WSRT) data of the Crab pulsar.

Left image courtesy of I. Joliot-Curie & F. Joliot/NMSI. Right image of SNR E0102 courtesy of NASA/CXC/STScI/MIT/SOA/D.Dewey/J.DePasquale; overlay of Crab pulsar radio data courtesy of J. van Leeuwen/ASTRON.

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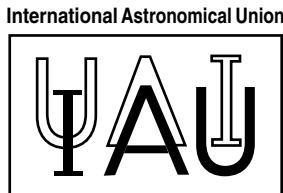
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INTERNATIONAL ASTRONOMICAL UNION  
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# NEUTRON STARS AND PULSARS: CHALLENGES AND OPPORTUNITIES AFTER 80 YEARS

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THE INTERNATIONAL ASTRONOMICAL UNION  
HELD IN BEIJING, CHINA  
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Edited by

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## Preface

The neutron hit the paraffin wax. It had been cast out when the alpha particle and beryllium nucleus had merged. In the wax, the neutron smashed out a proton. It traveled through the round bubble chamber, its mark captured on photographic plate. It's 1932, and the neutron is discovered.

A massive star exploded, and left a round supernova remnant. The neutron star was smashed out from the center of the crash. It spins and sends a beam of radio emission racing through the galaxy, to the telescope, arriving in 2012.

These two images, so visually similar as the front cover makes clear, mark the begin and current state of 80 years of neutron star research. That status, those newest results in neutron-star and pulsar studies, were presented at the IAUS291, Beijing, August 2012.

Several of the outstanding presentations in this volume are clearly linked to previous highlights in the lifetime of this octogenarian field. Chamel uses the sudden spin changes seen in some radio pulsars to constrain how the neutrons inside the star behave – 80 years after the initial idea by Landau that such stars might exist. And while Baade and Zwicky proposed in 1934 that neutron stars form in supernovae, Sumiyoshi now presents our, partial, understanding of the mechanism driving these explosions. The new LOFAR discoveries presented by Kondratiev, and by Coenen, were made using a telescope operating at the same frequencies, and build as a similar array of dipoles, as the original Cambridge array with which Hewish and Bell found the first radio pulsar in 1967. Backer, much missed, found the first, isolated millisecond pulsar in 1982 – a discovery that echoed in Roberts' review on the now numerous detections of black-widow millisecond pulsars.

Other results, however, are exceedingly novel. Saz Parkinson presented tens of gamma-ray pulsars that were blindly detected with Fermi-LAT, uncovering a new population that is nearby and energetic, and often radio-quiet. An entire session, headed off by Hobbs, showcased the potential for gravity-wave pulsar astronomy. In several talks the intriguing new-found relations between spin-down and profile evolution were discussed. Burke-Spolaor, and Karako, explained how some radio pulsars only emit sporadically.

On August 24, 2012, 17:12 Beijing time, this IAU Symposium 291 came to a close. Yet, the talks and posters remain, in several complementary forms: as both slides and video online†, and as proceedings in the volume before you.

*Joeri van Leeuwen  
21 November, 2012*

† <http://www.pulsarastronomy.net/IAUS291>

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## Welcome

The title of IAU Symposium 291 “Neutron Stars and Pulsars: Challenges and Opportunities after 80 Years” encapsulates the spirit of this Symposium: the 80 years since the idea of neutron stars was born and, in particular, the 45 years since pulsars were discovered have provided us with a rich harvest of scientific discovery, but many exciting avenues for future research remain. The Symposium, held in the huge Chinese National Convention Center adjacent to Beijing’s Olympic Park as part of the 28th General Assembly of the International Astronomical Union, was very successful with more than 160 talks and posters presented over the five days and 14 sessions of the meeting. We are pleased that most of these presentations are represented in these Proceedings.

The sessions covered current searches for pulsars, both radio and gamma-ray, neutron-star formation and properties, binary pulsars, pulsar timing and tests of gravitational theories, magnetars, radio transients, radio, X-ray and gamma-ray pulse properties and emission mechanisms, and future facilities. This range of topics illustrates the diverse nature and wide application of pulsar research. Exciting new results were presented in all sessions and it is impossible to list them all. However, I would like to mention the three plenary talks, presented by Scott Ransom, Nanda Ray and Michael Kramer, which were outstanding and given to a standing-room-only audience despite the early hour. As the corresponding articles in this Proceedings show, they managed to successfully communicate the excitement of current pulsar research. I had many comments afterward, mostly from “non-pulsar” people, about how fascinating these talks were. I would also like to give special mention to Jocelyn Bell-Burnell’s highly original closing remarks. We are grateful to her for allowing us to include them in this Proceedings.

Finally, I would like to give my thanks to the IAU and the GA Local Organising Committee for a well-run and successful meeting, to the Scientific Organising Committee for Symposium 291 for their assistance with putting together an excellent scientific programme, to all the presenters for realising the potential of the programme and to the Editor of the Proceedings, Joeri van Leeuwen, for all the hard work required to bring this volume to fruition. I hope and expect that it will be a valuable reference work for both current and future students and researchers in astronomy and astrophysics.

*R. N. (Dick) Manchester*

*16 November, 2012*