Abundances in LMC and SMC Globular Clusters

Jennifer A. Johnson

DAO/HIA/NRC, 5071 West Saanich Road, Victoria, BC, V9E 2E7, Canada

Inese I. Ivans

California Institute of Technology, MS 105-24,,Pasadena, CA, 91125, USA

Peter B. Stetson and James E. Hesser

DAO/HIA/NRC, 5071 West Saanich Road, Victoria, BC, V9E 2E7, Canada

Michael Bolte

UCO/Lick Observatory, University of California, Santa Cruz, CA 95064, USA

Abstract. We present results for abundances in four old Magellanic Cloud clusters based on high-resolution spectroscopy of individual giants.

1. Abundances

We present abundances for four old LMC and SMC clusters based on highresolution spectra of 2-3 red giants per cluster taken with the Magellan telescope. We find that in the two clusters close to the LMC bar, NGC 1898 and NGC 2019, the [Si/Fe] ratio is enhanced by ~ 0.5 dex relative to solar, while [Ca/Fe] and [Ti/Fe] are between 0 and 0.2 dex, similar to what is seen in the inner halo Milky Way clusters. In contrast, [Ca/Fe] is 0.3 dex in the outer LMC cluster Hodge 11 and and the SMC cluster NGC 121, which is the canonical value for the old outer halo Milky Way clusters. The [Fe/H] values found are in good agreement with those previously derived from the slopes of the red giant branches. Finally, the [Fe/H] and [α /Fe] ratios are compared to the results from integrated spectra of these clusters.