

EVOLUTION OF SUBSYSTEMS DURING COLLAPSE OF A CLUSTER

TOSHIO TSUCHIYA

Department of Astronomy, Kyoto University

Kyoto 606-01, Japan

Motivation: In the hierarchical clustering scenario, objects with smaller masses collapse and virialize earlier. I am interested in the effects of the existence of virialized subsystems on the structures of a collapsing larger mass object, and the fate of subsystems.

Comparison to previous works: Most work, both analytic and numerical, is concerned with the evolution of subsystems in a virialized cluster. Some numerical works have investigated dynamical effects in collapsing clusters, but these are usually not quite systematic.

Aim of this work: We try to study numerically the evolution of collapsing clusters with clumpy initial conditions. We make a systematic survey of two dimensional parameters: the number and the relative size of the subsystems.

Simulations: Numerical N-body simulations are made with the GRAPE-3A supercomputer. No cosmological expansion is included. Each subsystem is an equilibrium Plummer sphere. The initial radius of the cluster of the Plummer sphere is given by

$$R_0 \equiv 2r_v \left(\frac{M_{\text{total}}}{m} \right)^{(n_k+5)/6} = N_c^{(n_k+5)/6}, \quad (1)$$

where r_v is the initial radius of the cluster and the subsystems; M_{total} and m are the masses of the cluster and the subsystems and N_c is the number of the subsystems.

Results: We obtained a condition for the subsystems to survive the first violent phase of a collapsing cluster;

$$n_k \geq -1, \quad \text{and} \quad N_c \gg 1. \quad (2)$$