Recombination Line Abundances in NGC 6210

R. L. Kingsburgh¹, M. Komljenovic¹, J. A. López² and M. Peimbert²

¹York University, Dept. of Physics and Astronomy;

²Universidad Nacional Autónoma de México, Instituto de Astronomia

We present abundances derived from echelle spectroscopy of the planetary nebula NGC 6210 from the forbidden, collisionally excited lines, and from the permitted, recombination lines. We have obtained spectra from 3 positions in the nebula: at the central star, and 6 and 12 arcseconds south of the central star. We find a discrepancy of a factor of 2 between the O²⁺/H⁺ ratio derived by the recombination lines and by the forbidden lines in each spatial position. This discrepancy may result from the presence of temperature fluctuations in NGC 6210, where the collisionally excited lines are exponentially weighted to regions of higher temperature. A value for the root-mean-square temperature fluctuation parameter $t^2 \binom{OII}{abun} = 0.038 \binom{+0.007}{-0.015}$ for the 12 arcsecond offset spectrum. Thus no significant variation in t^2 is seen as a function of spatial position within the nebula. We have also estimated t^2 by comparing the derived Balmer Jump temperature at the 6 arcsecond offset, $T_e(BaJ) = 7900 \binom{+2400}{-1600}K$, with the temperature derived from the [O III] 5007/4363Å ratio $T_e[O III] = 9370 \pm 100K$ at this position. This comparison suggest $t^2(BaJ) = 0.041 \pm 0.020$, in agreement with the values for t^2 as estimated from the recombination lines.

Table 1 presents the O^{2+}/H^+ ratio derived for each transition of multiplet M at each spatial position. Table 1 also presents the average O^{2+}/H^+ ratio adopted for each position, the ratio of the O^{2+}/H^+ abundance derived via the recombination lines to that derived via the collisional lines $O^{2+}(\text{recom})/O^{2+}(\text{coll})$, the electron temperature derived via the [O III] 5007/4363Å line ratio t_e(O III) (in units of 10^4 K), and the implied t^2 based on the discrepancy between the recombination line and collisional line abundances.

$\lambda(\dot{A})$	M	transition	O ²⁺ /H ⁺ ×10 ³		
			CS posn	6" offset	12" offset
4072.16	10	$3d^4F_{5/2}-3p^4D_{3/2}$	1.15±0.08	0.845±0.17	1.22 ± 0.24
4153.30	19	$3d^4P_{5/2}-3p^4P_{3/2}$	1.12±0.17	0.784 ± 0.12	
4119.22	20	$3d^4D_{7/2}-3p^4P_{5/2}$	1.34 ± 0.13	0.932 ± 0.16	
4924.60	28	$3d^4P_{5/2}-3p^4S_{3/2}$	1.35 ± 0.20	0.853 ± 0.17	
4275.56	67	$4fF[4]_{9/2}-3d^4D_{7/2}$	1.49 ± 0.18	1.11 ± 0.22	
		Adopted	1.25 ± 0.15	$0.905 {\pm} 0.18$	1.22 ± 0.24
$O^{2+}(recom)/O^{2+}(coll)$		2.0	1.9	2.2	
		t _e (OIII)	0.900 ± 0.010	0.940 ± 0.010	0.940 ± 0.010
		$t^2(\underset{abun}{on})$	$0.038_{-0.0090}$	$0.040_{-0.0125}$	$0.051_{-0.015}$

Table 1